## CONSTRUCTION OF PERIODIC SOLUTIONS OF AUTONOMOUS SYSTEMS HITI ONE DEGREE OF FREEDOM IN THE CASE OF arBITRARY REAL ROOTS OF THE EQUATION FOR THE BASIC AMPLITUDES

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The method for the construction of periodic solutions of autonomous systems with one degree of freedom is worked out in detail for the case of simple nonzero roots of the equation for the basic amplitudes [1,2]. The present paper considers the general case when the roots of the abovementioned equation, being real and nonnegative numbers, can have any multiplicity.

1. Consider a nonlinear vibrating system with one degree of freedom

$$
\begin{equation*}
\frac{d^{2} x}{d t^{2}}+k^{2} x=\mu f\left(x, \frac{d x}{d t}, \mu\right) \tag{1.1}
\end{equation*}
$$

The function $f(x, d x / d t, \mu)$ will be considered analytic with respect to its arguments in a certain domain. The quantity $\mu$ is a small parameter.

Introduce a new independent variable $r=k t$. Then equation (1.1) assumes the form

$$
\begin{equation*}
\frac{d^{2} x}{d \tau^{2}}+x=\frac{\mu}{k^{2}} f\left(x, k \frac{d x}{d \tau}, \mu\right) \tag{1.2}
\end{equation*}
$$

In what follows the derivatives with respect to $r$ will be denoted by a letter with a prime.

Since the system is autonomous, then without loss of generality one of the initial conditions can be assumed to be

$$
\begin{equation*}
x^{\prime}(0)=0 \tag{1.3}
\end{equation*}
$$

Corresponding to this assumption the solution of the generating equation ( $\mu=0$ )

$$
x_{0}^{\prime \prime}(\tau)+x_{0}(\tau)=0
$$

has the form $x_{0}(r)=A_{0} \cos r$.
We shall seek periodic solutions of the basic equation (1.2) by the method of a small parameter. Let the second initial condition be of the form

$$
\begin{equation*}
x(0)=A_{0}+\beta \tag{1.4}
\end{equation*}
$$

where $\beta$ is a function of $\mu$ vanishing for $\mu=0$. Then the unknown function $x$ will have the form $x=x(r, \beta, \mu)$.

It is well-known that the period of vibrations of the autonomous system (1.2) depends on the parameter $\mu$ and may be put in the form of a sum $T=2 \pi+a$, where $2 \pi$ is the period of the generating solution and $a$ is a certain function of $\mu$ vanishing for $\mu=0$.

Let us determine the structure of the function $x(r, \beta, \mu)$. Assuming that this function can be developed in a series of integral powers of the parameters $\beta$ and $\mu$, we shall have

$$
x(\tau, \beta, \mu)=x_{0}(\tau)+B_{1}(\tau) \beta+C_{1}(\tau) \mu+B_{2}(\tau) \beta^{2}+D_{1}(\tau) \beta \mu+C_{8}(\tau) \mu^{2}+\ldots
$$

The functions $B_{n}(r)$ satisfy the equation

$$
B_{n}^{\prime \prime}(\tau)+B_{n}(\tau)=0 \quad(n=1,2, \ldots)
$$

the initial conditions being

$$
B_{1}(0)=1, \quad B_{1}^{\prime}(0)=0, \quad B_{n}(0)=0, \quad B_{n}^{\prime}(0)=0 \quad(n=2,3, \ldots)
$$

Consequently,

$$
B_{1}(\tau)=\cos \tau, \quad B_{n}(\tau)=0 \quad(n=2,3, \ldots)
$$

Taking this into account, the function $x(r, \beta, \mu)$ can be represented in the form

$$
\begin{equation*}
x(\tau, \beta, \mu)=A_{0} \cos \tau+\beta \cos \tau+\sum_{n=1}^{\infty}\left(C_{n}+\frac{\partial C_{n}}{\partial \beta} \beta+\frac{1}{2} \frac{\partial^{2} C_{n}}{\partial \beta^{2}} \beta^{2}+\ldots\right) \mu^{n} \tag{1.5}
\end{equation*}
$$

It is necessary to remember that all $C_{n}(r)$ and their derivatives with respect to $\beta$ are evaluated for $\beta=\mu=0$. From formula (1.5) it follows that the identity

$$
\begin{equation*}
\left(\frac{\partial^{m} x}{\partial ફ^{m}}\right)_{\mu=0}=\left(\frac{\partial^{m} x}{\partial A_{0}^{m}}\right)_{\mu=0} \tag{1.6}
\end{equation*}
$$

holds.

It is obvious that this property is possessed not only by $x$ but also by any of its derivatives with respect to $r$. In particular,

$$
\begin{equation*}
\left(\frac{\partial^{m} x^{\prime}}{\partial \beta^{m}}\right)_{\mu=0}=\left(\frac{\partial^{m} x^{\prime}}{\partial A_{0}^{m}}\right)_{\mu=0} \tag{1.7}
\end{equation*}
$$

Let us prove that differentiation with respect to $\beta$ can be replaced by differentiation with respect to $A_{0}$ also in mixed derivatives of $x$ with respect to $\beta$ and $\mu$, these derivatives being evaluated for the zero values of the parameters. First notice that under the assumptions made the function $f\left(x, k x^{\prime} ; \mu\right)$ can also be developed in a double series of integral powers of $\beta$ and $\mu$.

Consider the coefficient of $\beta^{n} \mu^{n+1}$ in the expansion of $x(r, \beta, \mu)$. This coefficient can be obtained by solving the corresponding nonhomogeneous differential equation of the second order for the initial conditions equal to zero. Multiplying both sides of this solution by $m!(n+1)$ ! we obtain

$$
\begin{equation*}
\left(\frac{\partial^{m+n+1} x}{\partial \beta^{m} \partial \mu^{n+1}}\right)_{0}=\frac{n+1}{k^{2}} \int_{0}^{\tau}\left(\frac{\partial^{m+n} f}{\partial \beta^{m} \partial \mu^{n}}\right)_{0} \sin \left(\tau-\tau_{1}\right) d \tau_{1} \tag{1.8}
\end{equation*}
$$

Assume that

$$
\begin{equation*}
\left(\frac{\partial^{m+n} x}{\partial \beta^{m} \partial \mu^{n}}\right)_{\beta=0, \mu=0}=\left(\frac{\partial^{m+n} x}{\partial A_{0}^{m} \partial \mu^{n}}\right)_{\beta=0, \mu=0} \tag{1.9}
\end{equation*}
$$

and that a similar identity exists for $x^{\prime}$. Let us prove that analogous formulas hold for the mixed derivative of the order $m+(n+1)$.

Since
then

$$
\frac{\partial^{m+n} f}{\partial \beta^{m} \partial \mu^{n}}=F_{m, n}\left(\frac{\partial x}{\partial \beta}, \frac{\partial x}{\partial \mu^{\mu}}, \ldots, \frac{\partial^{m+n} x}{\partial \beta^{m} \partial \mu^{n}}, \frac{\partial x^{\prime}}{\partial \beta}, \frac{\partial x^{\prime}}{\partial \mu}, \ldots, \frac{\partial^{m+n} x^{\prime}}{\partial \beta^{m} \partial \mu^{n}}\right)
$$

$$
\left(\frac{\partial^{m+n} f}{\partial \beta^{m} \partial \mu^{n}}\right)_{\beta=0, \mu=0}=\left(\frac{\partial^{m+n} f}{\partial A_{0}^{m} \partial \mu^{n}}\right)_{\beta=0, u=0}
$$

From the formula (1.8) and the formula obtained by differentiating both of its sides with respect to $r$, the sought relations immediately follow. The identity (1.9) holds for $n=0$, and, consequently, it will hold also for any $n$. Analogous identities hold for any derivatives of $x$ with respect to $r$.

Making use of these properties, the function $\boldsymbol{x}(r, \beta, \mu)$ can be represented in the form

$$
\begin{equation*}
x(\tau, \beta, \mu)=A_{0} \cos \tau+\beta \cos \tau+\sum_{n=1}^{\infty}\left(C_{n}+\frac{\partial C_{n}}{\partial A_{0}} \beta+\frac{1}{2} \frac{\partial^{2} C_{n}}{\partial A_{0}^{2}} \beta^{2}+\ldots\right) \mu^{n} \tag{1.10}
\end{equation*}
$$

Consequently, in order to construct the function $x(r, \beta, \mu)$, it is necessary first to know the coefficients $C_{n}(r)$ of $\mu^{n}$. The remaining coefficients of the double series can then be obtained by successive differentiation of $C_{n}(r)$ with respect to $A_{0}$. Denote

$$
H_{n}(\tau)=\frac{1}{(n-1)!}\left(\frac{d^{n-1} f}{d \mu^{n-1}}\right)_{\beta=0, \mu=0}
$$

where $d f / d \mu$ is the complete partial derivative of the function $f\left(x, k x^{\prime}, \mu\right)$ with respect to $\mu$. The coefficients $C_{n}(r)$ satisfy the equation

$$
C_{n}^{\prime \prime}(\tau)+C_{n}(\tau)=\frac{1}{k^{2}} H_{n}(\tau)
$$

the initial conditions being $C_{n}(0)=0$ and $C_{n}{ }^{\prime}(0)=0$. From this it follows that

$$
\begin{equation*}
C_{n}(\tau)=\frac{1}{k^{2}} \int_{0}^{\tau} H_{n}\left(\tau_{1}\right) \sin \left(\tau-\tau_{1}\right) d \tau_{1}, \quad C_{n}^{\prime}(\tau)=\frac{1}{k^{2}} \int_{0}^{\tau} H_{n}\left(\tau_{1}\right) \cos \left(\tau-\tau_{1}\right) d \tau_{1} \tag{1.11}
\end{equation*}
$$

The first four functions $H_{n}(r)$ are

$$
\begin{equation*}
H_{1}(\tau)=f\left(x_{0}, k x_{0}{ }^{\prime}, \mu\right)=f\left(A_{0} \cos \tau,-k A_{0} \sin \tau, 0\right) \tag{1.12}
\end{equation*}
$$

$$
\begin{equation*}
H_{2}(\tau)=\left(\frac{\partial f}{\partial x}\right)_{0} C_{1}+\left(\frac{\partial f}{\partial x^{\prime}}\right)_{0} C_{1}^{\prime}+\left(\frac{\partial f}{\partial \mu}\right)_{0} \tag{1.13}
\end{equation*}
$$

$$
H_{3}(\tau)=\frac{1}{2}\left(\frac{\partial^{2} f}{\partial x^{2}}\right)_{0} C_{1}^{2}+\frac{1}{2}\left(\frac{\partial^{2} f}{\partial x^{\prime 2}}\right)_{0} C_{1}^{\prime 2}+\frac{1}{2}\left(\frac{\partial^{2} f}{\partial \mu^{2}}\right)_{0}+\left(\frac{\partial^{2} f}{\partial x \partial x^{\prime}}\right)_{0} C_{1} C_{1}^{\prime}+
$$

$$
\begin{equation*}
+\left(\frac{\partial^{2} f}{\partial x \partial \mu}\right)_{0} C_{1}+\left(\frac{\partial^{2} f}{\partial x^{\prime} \partial \mu}\right)_{0} C_{1}^{\prime}+\left(\frac{\partial f}{\partial x}\right)_{0} C_{2}+\left(\frac{\partial f}{\partial x^{\prime}}\right)_{0} C_{2}^{\prime} \tag{1.14}
\end{equation*}
$$

$$
H_{4}(\tau)=\frac{1}{6}\left(\frac{\partial^{3} f}{\partial x^{3}}\right)_{0} C_{1}^{3}+\frac{1}{6}\left(\frac{\partial^{3} f}{\partial x^{\prime 3}}\right)_{0} C_{1}{ }^{3}+\frac{1}{6}\left(\frac{\partial^{3} f}{\partial \mu^{3}}\right)_{0}+\frac{1}{2}\left(\frac{\partial^{3} f}{\partial x^{2} \partial x^{\prime}}\right)_{0} C_{1}{ }^{2} C_{1}^{\prime}+
$$

$$
+\frac{1}{2}\left(\frac{\partial^{3} f}{\partial x \partial x^{\prime 2}}\right)_{0} C_{1} C_{1}^{\prime 2}+\frac{1}{2}\left(\frac{\partial^{3} f}{\partial x^{2} \partial \mu}\right)_{0} C_{1}{ }^{2}+\frac{1}{2}\left(\frac{\partial^{3} f}{\partial x^{\prime 2} \partial \mu}\right)_{0} C_{1}^{\prime 2}+\frac{1}{2}\left(\frac{\partial^{3} f}{\partial x \partial \mu^{2}}\right)_{0} C_{1}+
$$

$$
+\frac{1}{2}\left(\frac{\partial^{3} f}{\partial x^{\prime} \partial \mu^{2}}\right)_{0} C_{1}^{\prime}+\left(\frac{\partial^{3} f}{\partial x \partial x^{\prime} \partial \mu}\right)_{0} C_{1} C_{1}^{\prime}+\left(\frac{\partial^{2} f}{\partial x^{2}}\right)_{0} C_{1} C_{2}+\left(\frac{\partial^{2} f}{\partial x^{\prime 2}}\right)_{0} C_{1}^{\prime} C_{2}^{\prime}+
$$

$$
\begin{equation*}
+\left(\frac{\partial^{2} f}{\partial x \partial x^{\prime}}\right)_{0}\left(C_{1} C_{2}^{\prime}+C_{1}^{\prime} C_{2}\right)+\left(\frac{\partial^{2} f}{\partial x \partial \mu}\right)_{0} C_{2}+\left(\frac{\partial^{2} f}{\partial x^{\prime} \partial \mu}\right)_{0} C_{2}^{\prime}+\left(\frac{\partial f}{\partial x}\right)_{0} C_{3}+\left(\frac{\partial f}{\partial x^{\prime}}\right)_{0} C_{3}^{\prime} \tag{1.15}
\end{equation*}
$$

The subscript 0 indicates that in the partial derivatives of the function $f$ the variables $x, x^{\prime}$ and $\mu$ must be replaced by $A_{0} \cos \tau$, - $A_{0} \sin \tau$ and 0 respectively.
2. The periodicity conditions for the function $x(r, \beta, \mu)$ and its first derivative with respect to $r$, in conformity with the initial conditions (1.4) and (1.3), can be given in the form

$$
\begin{align*}
x(2 \pi+\alpha, \beta, \mu) & =A_{0}+\beta  \tag{2.1}\\
x^{\prime}(2 \pi+\alpha, \beta, \mu) & =0 \tag{2.2}
\end{align*}
$$

From condition (2.2) which defines the quantity a implicitly, let us determine $a=a(\beta, \mu)$ in the form of a double series of integral powers of $\beta$ and $\mu$. For the determination of the coefficients of this series it is necessary to calculate the partial derivatives of $a$ with respect to
$\beta$ and $\mu$ for $r=2 \pi$ and $\beta=\mu=0$.
The existence of partial derivatives of the implicit function $a(\beta, \mu)$ is assured in the present case by the condition

$$
\begin{equation*}
x^{\prime \prime}(2 \pi, 0,0)=-A_{0} \neq 0 \tag{2.3}
\end{equation*}
$$

Thus the construction of a series solution for the function $a(\beta, \mu)$ is possible only in the case when the amplitude of the generating solution is different from zero.

Since all derivatives of $x^{\prime}(r, \beta, \mu)$ with respect to $\beta$ are equal to zero for $r=2 \pi$ and $\mu=0$, then

$$
\left(\frac{\partial^{m} \alpha}{\partial \beta^{m}}\right)_{\mu=0}=0 \quad(m=1,2, \ldots)
$$

On the basis of the identity (1.9) in mixed derivatives the differentiation of the function $a$ with respect to $\beta$ can be replaced by differentiation with respect to $A_{0}$, i.e.

$$
\left(\frac{\partial^{m+n} \alpha}{\partial \beta^{m} \partial \mu^{n}}\right)_{\beta=0, \mu=0}=\left(\frac{\partial^{m+n} \alpha}{\partial A_{0}^{m} \partial \mu^{n}}\right)_{\beta=0, \mu=0}
$$

Consequently, it is only necessary to calculate the partial derivatives of the function $a$ with respect to $\mu$. Calculating successively the first four derivatives, we obtain

$$
\begin{align*}
& \left(\frac{\partial \alpha}{\partial \mu}\right)_{0}=\frac{1}{A_{0}} C_{1}{ }^{\prime}(2 \pi)=N_{1}(2 \pi)  \tag{2.4}\\
& \left(\frac{\partial^{2} \alpha}{\partial \mu^{2}}\right)_{0}=\frac{2}{A_{0}}\left[C_{2}{ }^{\prime}(2 \pi)+\frac{1}{k^{2}} H_{1}(2 \pi) N_{1}(2 \pi)\right]=2 N_{2}(2 \pi)  \tag{2.5}\\
& \left(\frac{\partial^{3} a}{\partial \mu^{3}}\right)_{0}=\frac{6}{A_{0}}\left\{C_{3}{ }^{\prime}(2 \pi)+\frac{1}{k^{2}} H_{1}(2 \pi) N_{2}(2 \pi)-\left[C_{2}(2 \pi)+\frac{1}{3 A_{0}} C_{1}{ }^{2}(2 \pi)-\right.\right. \\
& \left.\left.-\frac{1}{2 k^{2} A_{0}} H_{1}{ }^{\prime}(2 \pi) C_{1}{ }^{\prime}(2 \pi)-\frac{1}{k^{2}} H_{2}(2 \pi)\right] N_{1}(2 \pi)\right\}=6 N_{3}(2 \pi)  \tag{2.6}\\
& \left(\frac{\partial^{6} \alpha}{\partial \mu^{4}}\right)_{0}=\frac{24}{A_{0}}\left\{C_{4}^{\prime}(2 \pi)+\frac{1}{k^{2}} H_{1}(2 \pi) N_{3}(2 \pi)-\left[C_{2}(2 \pi)-\frac{1}{k^{2}} H_{2}(2 \pi)\right] N_{2}(2 \pi)-\right. \\
& -\left[C_{3}(2 \pi)+\frac{1}{A_{0}} C_{2}{ }^{\prime}(2 \pi) C_{1}{ }^{\prime}(2 \pi)-\frac{1}{k^{2} A_{0}} H_{1}{ }^{\prime}(2 \pi) C_{2}{ }^{\prime}(2 \pi)+\frac{2}{3 k^{2} A_{0}{ }^{2}} H_{1}(2 \pi) C_{1}{ }^{\prime 2}(2 \pi)-\right. \\
& -\frac{1}{6 k^{2} A_{0}{ }^{2}} H_{1}{ }^{\prime \prime}(2 \pi) C_{1}{ }^{\prime 2}(2 \pi)-\frac{1}{k^{4} A_{0}{ }^{2}} H_{1}(2 \pi) H_{1}{ }^{\prime}(2 \pi) C_{1}{ }^{\prime}(2 \pi)- \\
& \left.\left.-\frac{1}{2 k^{2} A_{0}} H_{2}{ }^{\prime}(2 \pi) C_{1}^{\prime}(2 \pi)-\frac{1}{k^{2}} H_{3}(2 \pi)\right] N_{1}(2 \pi)\right\}=24 N_{4}(2 \pi) \tag{2.7}
\end{align*}
$$

Hence for the function $a$ we obtain the expression

$$
\begin{equation*}
\alpha(\beta, \mu)=\sum_{n=1}^{\infty}\left(N_{n}+\frac{\partial N_{n}}{\partial A_{0}} \beta+\frac{1}{2} \frac{\partial^{2} N_{n}}{\partial A_{0}^{2}} \beta^{2}+\ldots\right) \mu^{n} \tag{2.8}
\end{equation*}
$$

Let us note that all quantities $N_{n}$ and their derivatives with respect to $A_{0}$ are evaluated for $r=2 \pi$ and $\hat{\beta}=\mu=0$.

Now consider the condition (2.1). For a preliminary simplification of
this condition expand the left-hand sides of (2.1) and (2.2) in series of powers of $a$. Multiply the equality (2.2) by $a$ and subtract it term-byterm from the equality (2.1). We then obtain

$$
\begin{gather*}
x(2 \pi, \beta, \mu)-\frac{1}{2} \alpha^{2} x^{\prime \prime}(2 \pi, \beta, \mu)-\frac{1}{3} \alpha^{3} x^{\prime \prime \prime}(2 \pi, \beta, \mu)-\frac{1}{8} \alpha^{4} \alpha^{\prime \prime \prime \prime}(2 \pi, \beta, \mu)-\ldots= \\
=A_{0}+\beta \tag{2.9}
\end{gather*}
$$

Substitute in this equality the expressions for $x(2 \pi, \beta, \mu), x^{\prime \prime}(2 \pi$, $\beta, \mu$ ), ... as given by formula (1.10) and for $a$ its expression by formula (2.8). The terms of the left-hand side of this equality are products of two functions, each of which possesses the property (1.9). Terms on the same side which are independent of $\mu$ are compensated by the right-hand side. Thus the condition (2.9) can be put in the form

$$
\begin{equation*}
\sum_{n=1}^{\infty}\left(M_{n}+\frac{\partial M_{n}}{\partial A_{0}} \beta+\frac{1}{2} \frac{\partial^{2} M_{n}}{\partial A_{0}^{2}} \beta^{2}+\ldots\right) \mu^{n}=0 \tag{2.10}
\end{equation*}
$$

Let us note that the quantities $M_{n}$ are also evaluated for $r=2 \pi$ and $\beta=\mu=0$ : Calculating them we obtain

$$
\begin{gather*}
M_{1}(2 \pi)=C_{1}(2 \pi)  \tag{2.11}\\
M_{2}(2 \pi)=C_{2}(2 \pi)+\frac{1}{2 A_{0}} C_{1}^{\prime 2}(2 \pi)  \tag{2.12}\\
M_{3}(2 \pi)=C_{3}(2 \pi)+\frac{1}{A_{0}} \cdot\left[C_{2}^{\prime}(2 \pi)+\frac{1}{2 k^{2} A_{0}} H_{1}(2 \pi) C_{1}^{\prime}(2 \pi)\right] C_{1}^{\prime}(2 \pi)  \tag{2.13}\\
M_{4}(2 \pi)=C_{4}(2 \pi)+\frac{1}{2 A_{0}} \cdot C_{2}^{\prime 2}(2 \pi)+\frac{1}{A_{0}}\left[C_{3}^{\prime}(2 \pi)+\frac{1}{k^{2} A_{0}} H_{1}(2 \pi) C_{2}^{\prime}(2 \pi)-\right. \\
-\frac{1}{2 A_{0}} C_{2}(2 \pi) C_{1}^{\prime}(2 \pi)-\frac{1}{8 A_{0}^{2}} C_{1}^{\prime 3}(2 \pi)+\frac{1}{6 k^{2} A^{2}} H_{1}^{\prime}(2 \pi) C_{1}^{\prime 2}(2 \pi)+ \\
\left.+\frac{1}{2 k^{2} A_{0}} H_{2}(2 \pi) C_{1}^{\prime}(2 \pi)+\frac{1}{2 k^{4} A_{0}^{2}} H_{1}{ }^{2}(2 \pi) C_{1}^{\prime}(2 \pi)\right] C_{1}^{\prime}(2 \pi) \tag{2.14}
\end{gather*}
$$

3. The obtained formulas permit us to pass to the determination of the quantities $a$ and $\beta$ and to construct periodic solutions of the equation (1.2).

Assume that the quantity $\beta$ can be expanded in a power series of $\mu$, i.e.

$$
\begin{equation*}
\beta=\sum_{n=1}^{\infty} A_{n} \mu^{n} \tag{3.1}
\end{equation*}
$$

Substituting this series into the formula (2.10) and equating to zero the coefficients of all powers of $\mu$, we obtain*

$$
\begin{align*}
& M_{1}(2 \pi)=C_{1}(2 \pi)=0  \tag{3.2}\\
& A_{1} \frac{\partial C_{1}}{\partial A_{0}}+M_{2}(2 \pi)=0 \tag{3.3}
\end{align*}
$$

[^0]\[

$$
\begin{gather*}
A_{2} \frac{\partial C_{1}}{\partial A_{0}}+\frac{1}{2} A_{1}^{2} \frac{\partial^{2} C_{1}}{\partial A_{0}^{2}}+A_{1} \frac{\partial M_{2}}{\partial A_{0}}+M_{3}(2 \pi)=0  \tag{3.4}\\
A_{3} \frac{\partial C_{1}}{\partial A_{0}}+A_{2}\left(\frac{\partial M_{2}}{\partial A_{0}}+A_{1} \frac{\partial^{2} C_{1}}{\partial A_{0}^{2}}\right)+\frac{1}{6} A_{1}^{3} \frac{\partial^{3} C_{2}}{\partial A_{0}^{8}}+ \\
\quad+\frac{1}{2} A_{1}^{2} \frac{\partial^{2} M_{2}}{\partial A_{0}^{2}}+A_{1} \frac{\partial M_{8}}{\partial A_{0}}+M_{4}(2 \pi)=0 \tag{3.5}
\end{gather*}
$$
\]

From the obtained condition the amplitude $A_{0}$ of the generating solution and the coefficients $A_{n}(n=1,2, \ldots)$ determining the initial displacenent $\beta$ of the system can be calculated in succession.

The condition (3.2) is an amplitudinal equation from which $A_{0}$ can be calculated. If $A_{0}$ is not a multiple root of the equation (3.2), then from the remaining conditions, which are linear with respect to $A_{1}, A_{2}, \ldots$, these coefficients can be calculated in succession.

If equation (3.2) has a double root, then for such a root $\partial C_{1} / \partial A_{0}=0$.
In order that a periodic solution exist in the last case, it is necessary that the supplementary condition $M_{2}(2 \pi)=0$ be satisfied. This last fact is easily seen from the condition (3.3).

If this condition, together with that of (3.2), is identically satisfied, then $\partial M_{2} / \partial A_{0}=0$. The coefficient $A_{1}$ is given by the quadratic equation

$$
\frac{1}{2} A_{1}^{2}{\frac{\partial}{}{ }^{2} C_{1}}_{\partial A_{0}^{2}}{ }^{2}+M_{3}(2 \pi)=0
$$

Since the quantity $A_{1}$ must be real, then there can exist for $A_{1}$ either two values or none at all. The equation for the determination of the remaining coefficients $A_{2}, A_{3}, \ldots$ will again be linear.

The case of a triple root of the amplitude equation (3.2) can be discussed in a similar way. Here, for the existence of a periodic solution, it is necessary to satisfy one additional condition $M_{3}(2 \pi)=0$. The coefficient $A_{1}$ in such a case is determined by an equation of degree three.

It is obvious that for a root of multiplicity $n$ of the equation (3.2), $n-1$ supplementary conditions will have to be satisfied, and the coefficient $A_{1}$ will be given by an equation of degree $n$.

Thus, in the case of multiple roots of the equation for the basic amplitudes, bifurcation of the generating solution is possible.

In order to determine the period of the solution of the equation (1.2) assume that this period can be represented in the form of a series in powers of $\mu$,

$$
\begin{equation*}
T=2 \pi\left(1+h_{1} \mu+h_{2} \mu^{2}+\cdots\right) \tag{3.6}
\end{equation*}
$$

Then

$$
\begin{equation*}
\alpha=2 \pi \sum_{n=1}^{\infty} h_{n} \mu^{n} \tag{3.7}
\end{equation*}
$$

Substitute in the right-hand side of the formula (2.8) the expression (3.1) for $\beta$ and replace the left-hand side by the series (3.7). Equating the coefficients of equal powers of $\mu$ on both sides of the equality we obtain*

$$
\begin{gather*}
h_{1}=\frac{1}{2 \pi} N_{1}  \tag{3.8}\\
h_{2}=\frac{1}{2 \pi}\left(A_{1} \frac{\partial N_{1}}{\partial A_{0}}+N_{2}\right)  \tag{3.9}\\
h_{3}=\frac{1}{2 \pi}\left(A_{2} \frac{\partial N_{1}}{\partial A_{0}}+\frac{1}{2}{\left.A_{1}{ }^{2} \frac{\partial^{2} N_{1}}{\partial A_{0}{ }^{2}}+A_{1} \frac{\partial N_{2}}{\partial A_{0}}+N_{3}\right)}_{h_{4}=\frac{1}{2 \pi}\left[A_{3} \frac{\partial N_{1}}{\partial A_{0}}+A_{2}^{\prime}\left(\frac{\partial N_{2}}{\partial A_{0}}+A_{1} \frac{\partial^{2} N_{1}}{\partial A_{0}^{2}}\right)+\frac{1}{6} A_{1}{ }^{3} \frac{\partial^{3} N_{1}}{\partial A_{0}{ }^{3}}+\right.}^{\left.+\frac{1}{2} A_{1}{ }^{2} \frac{\partial^{2} N_{2}}{\partial A_{0}^{2}}+A_{1} \frac{\partial N_{3}}{\partial A_{0}}+N_{4}\right]}\right. \text {, } \tag{3.10}
\end{gather*}
$$

Let us make an additional substitution for the independent variable by putting

$$
\begin{equation*}
\tau=\tau_{0}\left(1+h_{1} \mu+h_{2} \mu^{2}+\cdots\right) \tag{3.12}
\end{equation*}
$$

and let us look for a solution of the equation (1.2) in terms of $r_{0}$. This solution has period $2 \pi$, independent of $\mu$. Substitute (3.12) for $r$ into the functions $C_{n}(r)$ and $\cos r$ and expand them in series of powers of $\mu$. We then obtain

$$
\begin{equation*}
r_{n}(-)=C_{n}\left(\tau_{0}\right)+h_{1} \tau_{0} C_{n}^{\prime}\left(\tau_{0}\right) \mu+\left[h_{2} \sigma_{0} C_{n}^{\prime}\left(\tau_{0}\right)+\frac{1}{2} h_{1}^{2} \tau_{0}^{2} C_{n}^{\prime \prime}\left(\tau_{0}\right)\right] \mu^{2}-\cdots \tag{3.13}
\end{equation*}
$$

and

$$
\begin{gather*}
\cos \tau=\cos \tau_{0}-h_{1} \tau_{10} \sin \tau_{0} \mu-\left(h_{2} \tau_{0} \sin \tau_{0}+\frac{1}{2} h_{1}^{2} \tau_{0}^{2} \cos \tau_{0}\right) \mu^{2}- \\
-\left(h_{3} \tau_{0} \sin \tau_{0}+h_{1} h_{2} \tau_{0}^{2} \cos \tau_{0}-\frac{1}{6} h_{1}^{3} \tau_{0}^{3} \sin \tau_{0}\right) \mu^{3}-\cdots \tag{3.14}
\end{gather*}
$$

Substitute into the formula (1.10) the expressions (3.13) and (3.14) for the functions $C_{n}(r)$ and cos $r$ and drop the subscript 0 of $r_{0}$. Expand the left-hand side of this formula in the form of a series in powers of the parameter $\mu$, i.e. let

[^1]\[

$$
\begin{equation*}
x(\tau)=x_{0}(\tau)+\mu x_{1}(\tau)+\mu^{2} x_{2}(\tau)+\cdots \tag{3.15}
\end{equation*}
$$

\]

Equating the coefficients of equal powers of $\mu$ on both sides of the equality, we obtain

$$
\begin{gather*}
x_{0}(\tau)=A_{0} \cos \tau  \tag{3.16}\\
x_{1}(\tau)=A_{1} \cos \tau+C_{1}(\tau)-h_{1} A_{0} \tau \sin \tau  \tag{3.17}\\
x_{2}(\tau)=A_{2} \cos \tau+C_{2}(\tau)+h_{1} \tau C_{1}^{\prime}(\tau)+A_{1} \frac{\partial C_{1}}{\partial A_{0}}- \\
-h_{1} A_{1} \tau \sin \tau-\frac{1}{2} h_{1}{ }^{2} A_{0} \tau^{2} \cos \tau-h_{2} A_{0} \tau \sin \tau  \tag{3.18}\\
x_{3}(\tau)=A_{3} \cos \tau+C_{3}(\tau)+h_{1} \tau C_{2}^{\prime}(\tau)+A_{1} \frac{\partial C_{2}}{\partial A_{0}}+h_{1} A_{1} \tau \frac{\partial C_{2}}{\partial A_{0}}+\frac{1}{2} h_{1}{ }^{2} \tau^{2} C_{1}^{\prime \prime}(\tau)+ \\
+h_{2} \tau C_{1}^{\prime}(\tau)+A_{2} \frac{\partial C_{1}}{\partial A_{0}}+\frac{1}{2} A_{1}{ }^{2} \frac{\partial^{2} C_{1}}{\partial A_{0}{ }^{2}}-h_{1} A_{2} \tau \sin \tau-\frac{1}{2}-h_{1}{ }^{2} A_{1} \tau^{2} \cos \tau- \\
-h_{2} A_{1} \tau \sin \tau+\frac{1}{6} h_{1}{ }^{3} A_{0} \tau^{3} \sin \tau-h_{1} h_{2} A_{0} \tau^{2} \cos \tau-h_{3} A_{0} \tau \sin \tau \tag{3.19}
\end{gather*}
$$

It is obvious that the functions $C_{n}(r)$ contain nonperiodic terms which are compensated by the terms depending on $h_{1}, h_{2}, \ldots, h_{n}$.
4. Consider the case of a zero root $A_{0}=0$ of the equation (3.2) for the basic amplitudes.

Let us note that one root of the amplitude equation is always zero, which may also be a multiple root of odd order. The preceding analysis, however, is not applicable in the case of a zero root.

By means of formula (3.12) introduce into equation (1.2) a new independent variable. Then we obtain

$$
\begin{equation*}
x^{\prime \prime}+h^{2} x=\mu \frac{h^{2}}{k^{2}} f\left(x, \frac{k}{h} x^{\prime}, \mu\right) \tag{4.1}
\end{equation*}
$$

where

$$
\begin{equation*}
h==1+h_{1} \mu+h_{2} \mu^{2}+\cdots \tag{4.2}
\end{equation*}
$$

Let us look for a solution of (4.1) in the form of a series (3.15). The function $x_{1}$ is determined by the equation

$$
x_{1}^{n}+x_{1}=k^{-2} f(0,0,0)
$$

The solution of this equation under the conditions (1.3) and (1.4) have the form

$$
x_{1}=P_{1} \cos \tau+k^{-2} f(0,0,0), \quad P_{1}=A_{1}-k^{-2} f(0,0,0)
$$

For the function $x_{2}$ we have the equation

$$
x_{2}^{\prime}+x_{2}=-2 h_{1} x_{1}+\frac{2 h_{1}}{k^{2}} f(0,0,0)+\frac{1}{k^{2}}\left[\left(\frac{\partial f}{\partial x}\right)_{0} x_{1}+\left(\frac{\partial f}{\partial x^{\prime}}\right)_{0} x_{1}^{\prime}+\left(\frac{\partial f}{\partial \mu}\right)_{0}\right]
$$

If $A_{0}=0$ is a simple root, then from the periodicity condition for $x_{2}$ it follows that $P_{1}=0$. In the case of a multiple zero root we have the condition

$$
\left[2 h_{1}-\frac{1}{k^{2}}\binom{\partial f}{\partial x}_{0}\right] P_{1}=0
$$

Under the assumption that the expression in the squared brackets is equal to zero, it follows from the periodicity condition for the function $x_{3}$ that $P_{1}=0$. Hence we have that $x_{1}=A_{1}$. From the conditions of periodicity for the immediately following functions $x_{n}$ we obtain that $x_{2}=A_{2}, x_{3}=A_{3}$, and so un. Consequently,

$$
\begin{equation*}
x=\sum_{n=1}^{\infty} A_{n} \mu^{n}=\beta(\mu) \tag{4.3}
\end{equation*}
$$

Thus the zero solution of the generating equation corresponds to the time independent solution of the complete equation.

Substitute the obtained value of $x$ in equation (1.2). Then

$$
\begin{equation*}
\beta=\mu k^{-2} f(\beta, 0, \mu) \tag{4.4}
\end{equation*}
$$

Expanding both sides of this equality in series of powers of $\mu$ and equating coefficients of equal power of $\mu$, we will obtain formulas for the successive determination of the coefficients $A_{n}$.

The case of the zero root corresponds to the equilibrium of the generating system. Consequently, the equilibrium of the generating solution goes over into the equilibrium of the complete system. Only the coordinate of the equilibrium position can change.
5. Let the function $f\left(x, k x^{\prime}, \mu\right)$ be independent of $\mu$ and have the form $f\left(x, k x^{\prime}\right)=k f_{1}(x) x^{\prime}$. In this case the expressions for $H_{x}(r)$ simplify to

$$
\begin{gather*}
H_{2}(\tau)=k \frac{\partial}{\partial \tau}\left[f_{1}\left(x_{0}\right) C_{1}\right]  \tag{5.1}\\
H_{3}(\tau)=k-\frac{\partial}{\partial \tau}\left[\frac{1}{2} f_{1}^{\prime}\left(x_{0}\right) C_{1}^{2}+f_{1}\left(x_{0}\right) C_{2}\right]  \tag{5.2}\\
H_{4}(\tau)=k \frac{\partial}{\partial \tau}\left[\frac{1}{6} f_{1}^{\prime \prime}\left(x_{0}\right) C_{1}^{3}+f_{1}^{\prime}\left(x_{0}\right) C_{1} C_{2}+f_{1}\left(x_{0}\right) C_{3}\right] \tag{5.3}
\end{gather*}
$$

It can be shown that in this case $C_{1}^{\prime}(2 \pi)=0$ and $C_{2}(2 \pi)=0$.
On the basis of (2.12) and (3.3) it follows that $A_{1}=0$ in the case of simple roots of the equation for the basic amplitudes. Further it can be proved that in the case of a root of multiplicity $n$ the equality $\partial^{n-1} C_{2} / \partial A_{0}{ }^{n-1}=0$ holds.
6. Finally, consider some examples. 1. Let the function $f\left(x, k x^{\prime}\right)$ be of the form

$$
\begin{equation*}
f\left(x, k x^{\prime}\right)=k\left(\alpha+\varepsilon x+\beta x^{2}+x x^{3}+\gamma x^{4}\right) x^{\prime} \tag{6.1}
\end{equation*}
$$

The equation for the basic amplitudes, after rejection of the zero root, is

$$
\gamma A_{0}{ }^{4}+2 \beta A_{0}{ }^{2}+8 \alpha=0
$$

The double root of this equation is

$$
A_{0}{ }^{2}=-\beta / \gamma, \quad \beta^{2}=8 \alpha \gamma
$$

The equation for the determination of the coefficient $A_{1}$ is

$$
A_{1}{ }^{2}+\frac{1}{576} A_{0}^{4} k^{-2}\left(\frac{1}{128} \gamma^{2} A_{0}^{6}+\frac{257}{100} x^{2} A_{0}^{4}+\frac{41}{5} \varepsilon x A_{0}^{2}+\frac{13}{2} \varepsilon^{2}\right)=0
$$

This equation does not possess real roots either in the case $\epsilon=\kappa=0$ or in the general case. Thus, in the case of multiple roots of the amplitude equation, periodic solutions of the equation (1.1) do not exist if the function $f\left(x, k x^{\prime}\right)$ is given by the formula (6.1).
2. Let $f\left(x, k x^{\prime}\right)$ be of the form

$$
\begin{equation*}
f\left(x, k x^{\prime}\right)=k\left(\alpha+\beta x^{2}+\gamma x^{4}+\delta x^{6}\right) x^{\prime} \tag{6.2}
\end{equation*}
$$

The equation for the basic amplitudes (besides the zero root) is

$$
5 \delta A_{0}{ }^{6}+8 \gamma A_{0}^{4}+16 \beta A_{0}^{2}+64 \alpha=0
$$

If one considers a double root of this equation, then an additional equation

$$
15 \delta A_{0}{ }^{4}+16 \gamma A_{0}{ }^{2}+16 \beta=0
$$

must be considered.
Evaluating $C_{1}(r)$ we obtain that

$$
\begin{aligned}
C_{1}(\tau) & =\frac{1}{64} A_{0}{ }^{5} k^{-1}\left[\frac{1}{48} \delta A_{0}{ }^{2} \sin 7 \tau+\left(\frac{5}{24} \delta A_{0}{ }^{2}+\frac{1}{6} \gamma\right) \sin 5 \tau-\right. \\
& \left.-\left(\frac{3}{4} \delta A_{0}{ }^{2}+\frac{1}{2} \gamma\right) \sin 3 \tau+\left(\frac{17}{16} \delta A_{0}{ }^{2}+\frac{2}{3} \gamma\right) \sin \tau\right]
\end{aligned}
$$

For the coefficient $A_{1}$ we have the equation

$$
\begin{gathered}
\left(15 \delta A_{0}{ }^{2}+8 \gamma\right) A_{1}{ }^{2}+\frac{1}{128^{2}} A_{0}{ }^{10} k^{-2}\left(\frac{2117}{354} \delta^{3} A_{0}{ }^{6}+\frac{1621}{144} \gamma^{2} A_{0}{ }^{4}+\right. \\
\left.+\frac{1189}{18} \gamma^{2} \delta A_{0}{ }^{2}+\frac{16}{9} \gamma^{3}\right)=0
\end{gathered}
$$

For $\delta=0$ a periodic solution will not exist (as a consequence of Example 1.). For $\delta \neq 0$ a sufficiently narrow interval will exist in which periodic vibrations are possible. These intervals are bounded by the zeros of the polynomials, being the coefficient of $A_{1}{ }^{2}$ and the free term of the equation. We have

$$
-1.875 A_{0}^{2}<\gamma / \delta<-1.527 A_{0}^{2}
$$

The upper bound is determined only approximately. The period of vibrations up to the terms containing $\mu^{3}$ is

$$
T=\frac{2 \pi}{k} \cdot\left[1+\frac{1}{64^{4}} A_{0}{ }^{8} k^{-2}\left(\frac{89}{32} \delta^{2} A_{0}^{4}+\frac{23}{6} \delta \gamma A_{0}{ }^{2}+\frac{4}{3} \gamma^{2}\right) \mu^{2}+\cdots\right]
$$

The vibrations themselves up to terms containing $\mu^{2}$ are given by the formula

$$
x(\tau)=A_{0} \cos \tau+\left[A_{1} \cos \tau+C_{1}(\tau)\right] \mu+\cdots
$$

where $A_{0}, A_{1}$ and $C_{1}(r)$ are determined by the above mentioned formulas.

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[^0]:    - The formulas (3.2) to (3.4) coincide with the periodicity conditions (21). (26) and (31) for the functions $a_{n}(r)$ in the paper [2].

[^1]:    - Formulas (3.8) to (3.10) coincide with the formulas (22), (27) and (32) of the paper [2], obtained from the conditions of periodicity for the functions $b_{n}(r)$.

